

# Telescopes



## History and Basics

By Steve Clark  
Galway Public Library Astronomy Club

# History



Hans Lippershey (1570 – 1719),  
1608: invented telescope

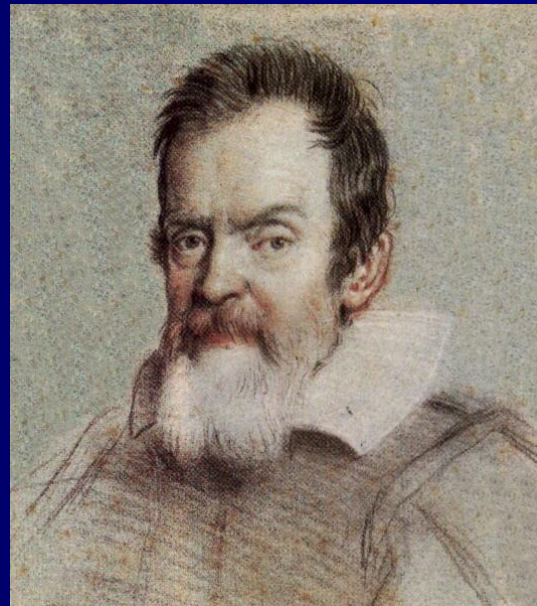


Credits to: Microsoft Encarta Encyclopedia 99 Deluxe

# History



Galileo Galilei, 1609 - 1632: 1<sup>st</sup> astronomical use of a telescope (Jupiter).



# History



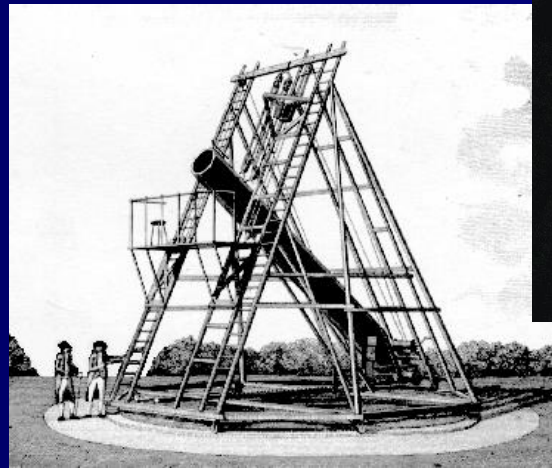
Isaac Newton, 1671: Newtonian Telescope  
Born to combat chromatic aberration.



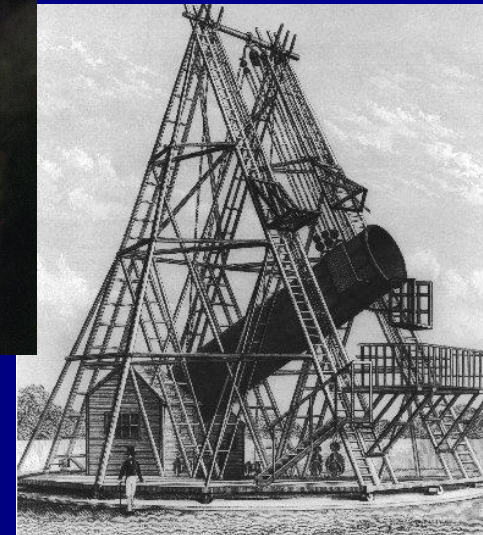
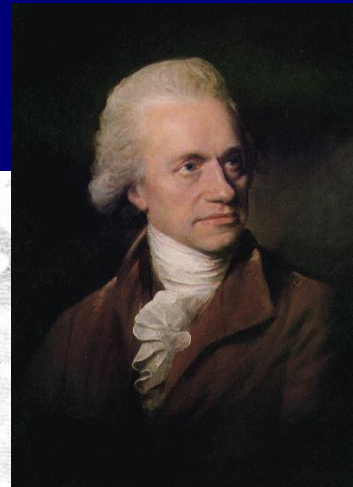
# History



William Herschel, 1780s: Large Reflectors



20 foot



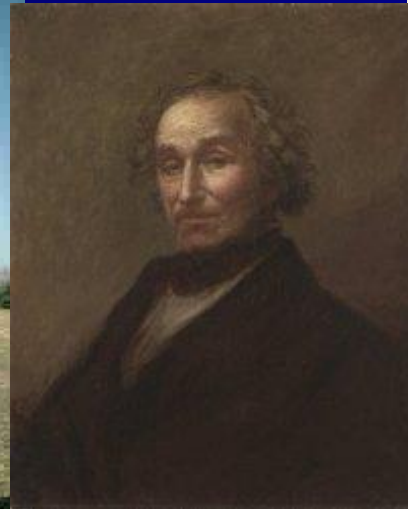
40 foot



# History



- Alvan Clark & Sons, 1880s: Large refractors – 40" Yerkes.



# History



- 60" and 100" Telescopes at Mt. Wilson – Beginning of the Big Reflectors, 1<sup>st</sup> use of steel construction.



# History



200" Hale at Mt. Palomar – Modern truss design,  
BIG construction era.

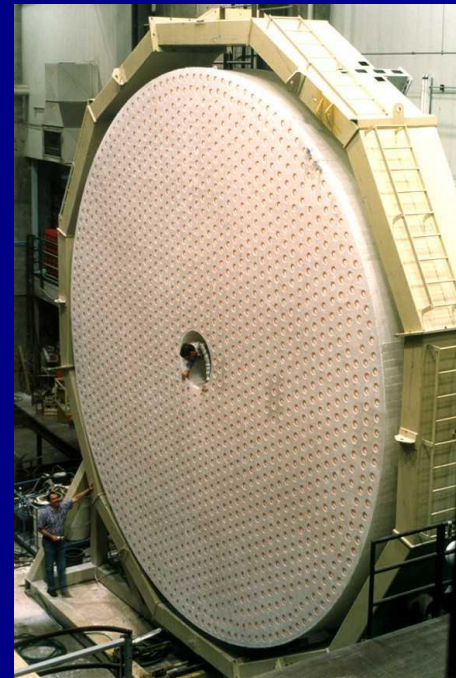




# History



Spin Casting up to 8.4m mirrors.



# History



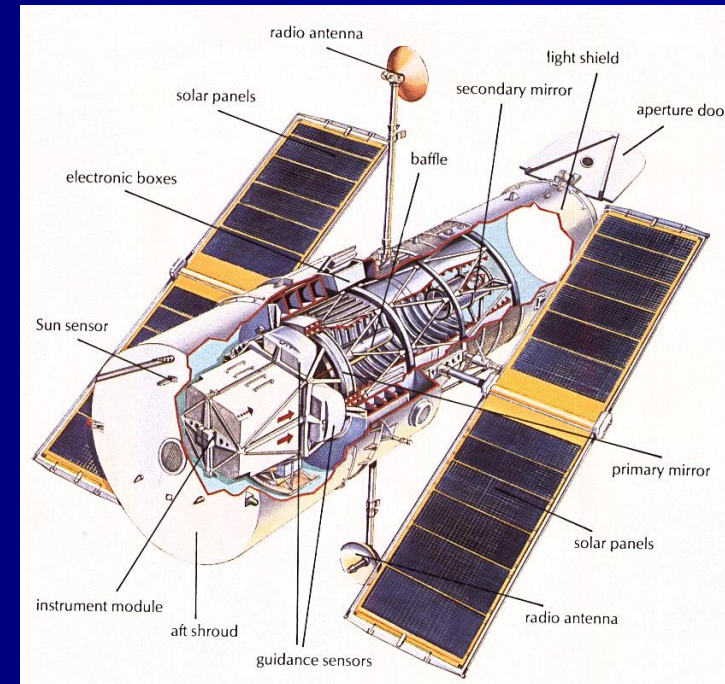
Twin 10m Kecks at Mauna Kea – Multiple mirror Technology.



# History



## ➤ Hubble Space Telescope.





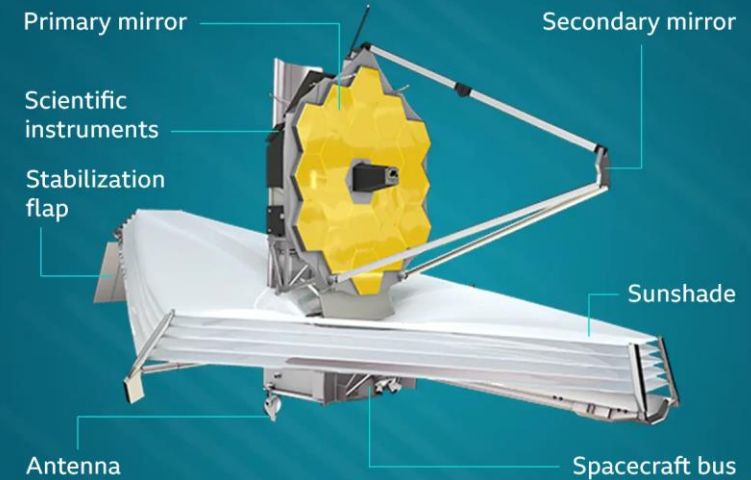
# History



## James Webb Space Telescope



### James Webb Space Telescope



Source: Nasa

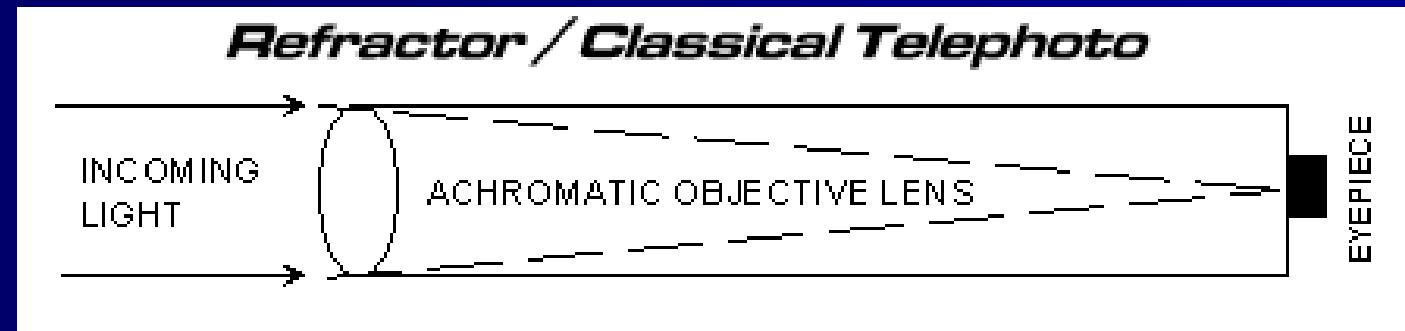




# Telescope Types



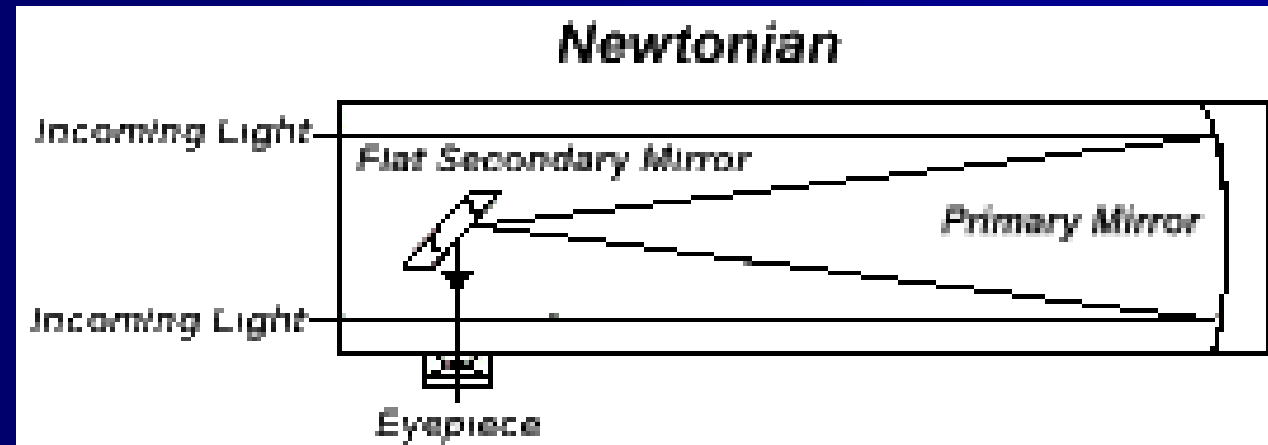
## Refractor



# Telescope Types



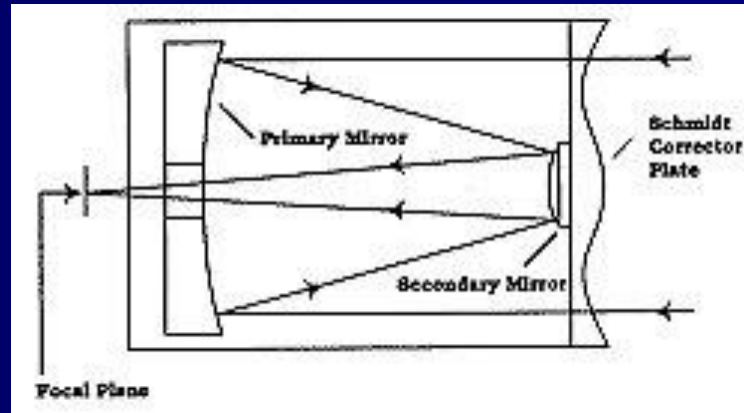
## Newtonian Reflector



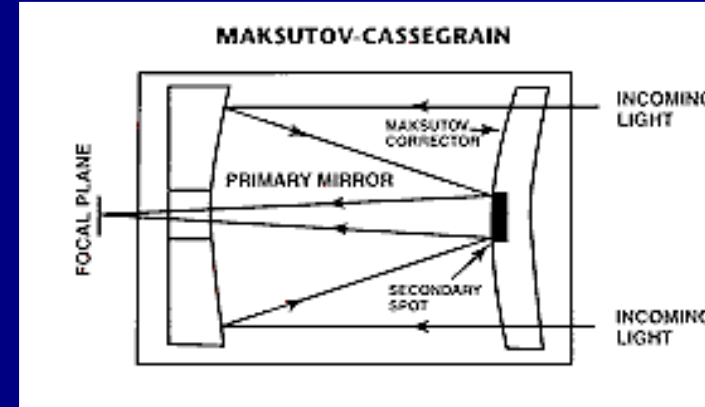
# Telescope Types



## Catadioptrics (CATs)

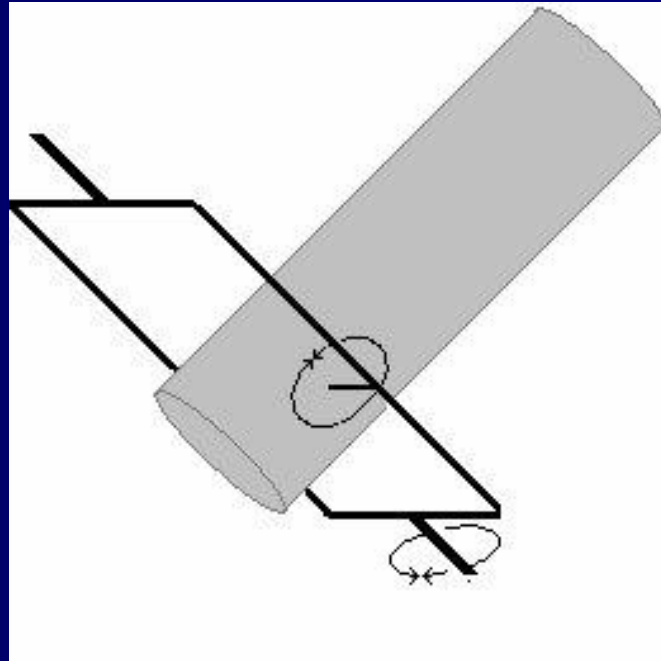


Schmidt-Cassegrain (SCT)

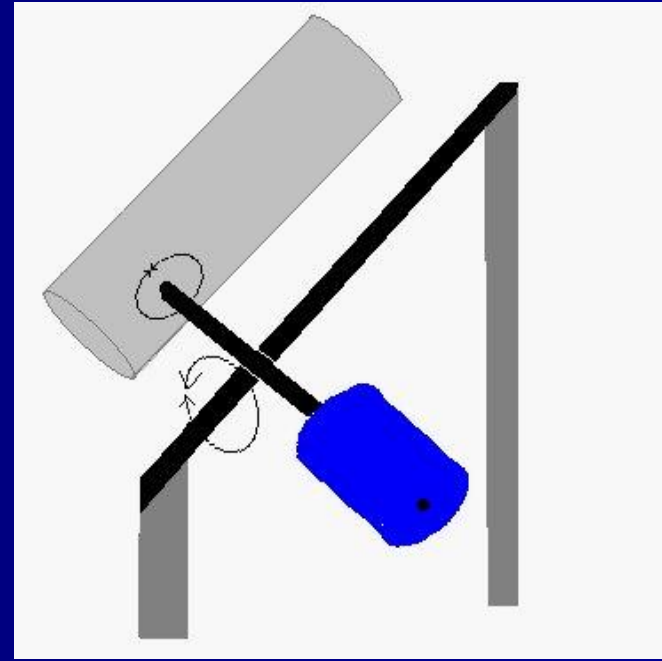


Maksutov-Cassegrain (Mak)

# Telescope Mounts



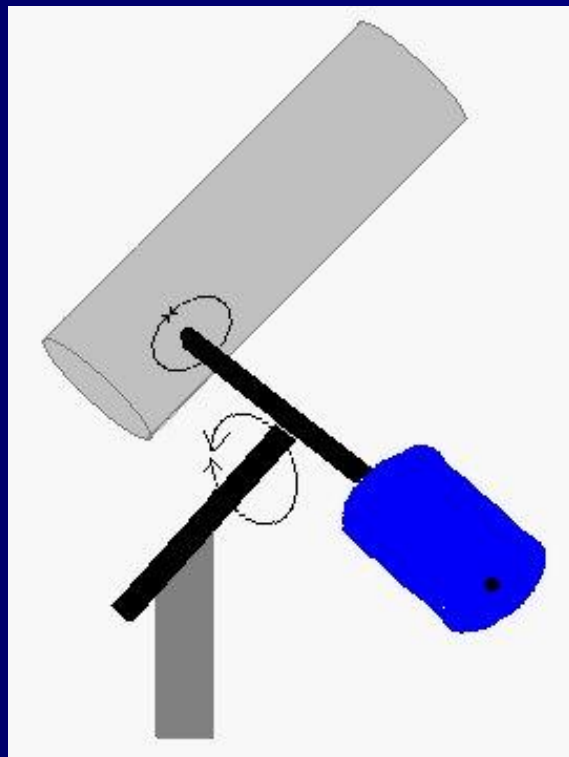
English Cradle



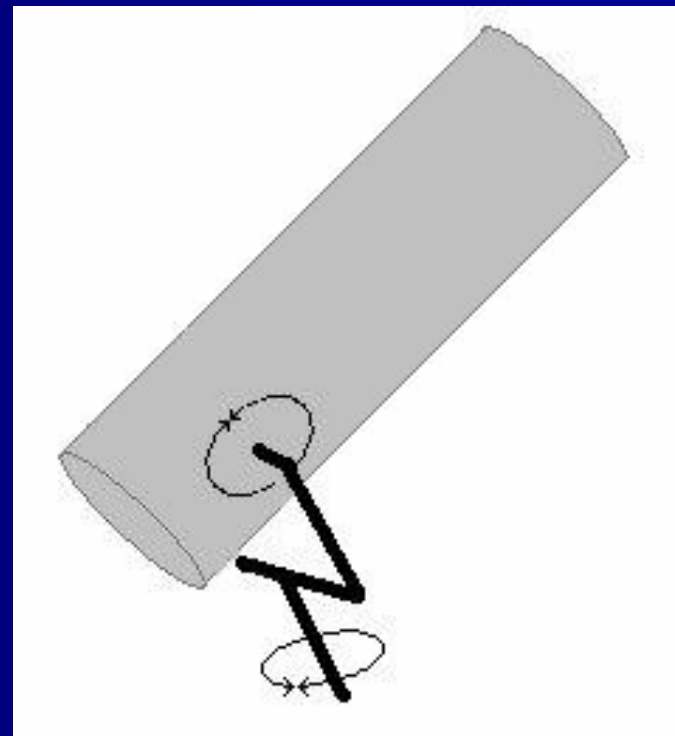
Cross Axis



# Telescope Mounts

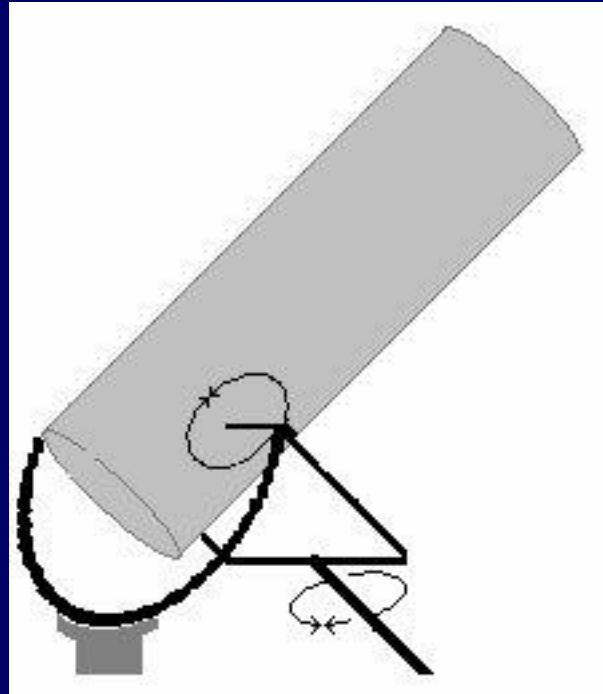


German Equatorial

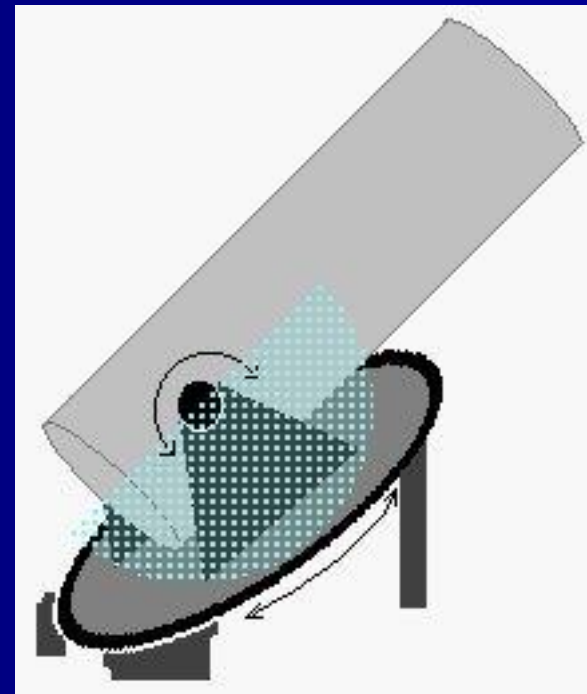


Fork Mount

# Telescope Mounts

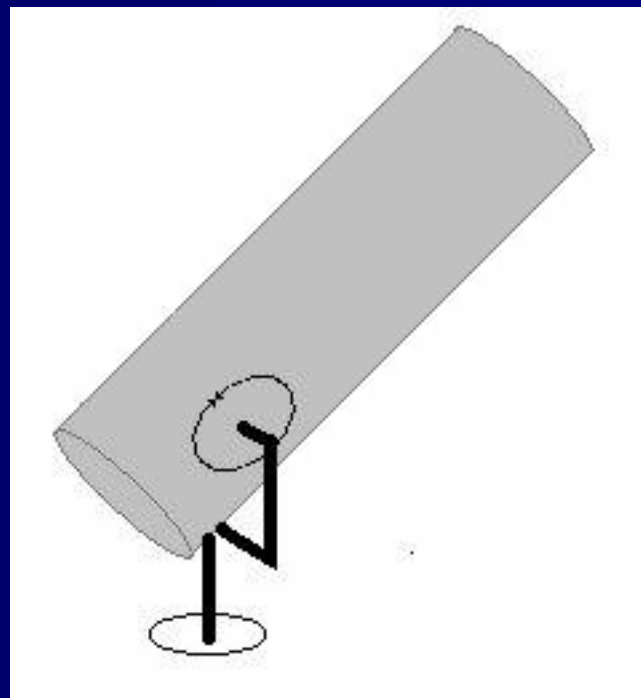


Horseshoe Mount



Polar Disk

# Telescope Mounts

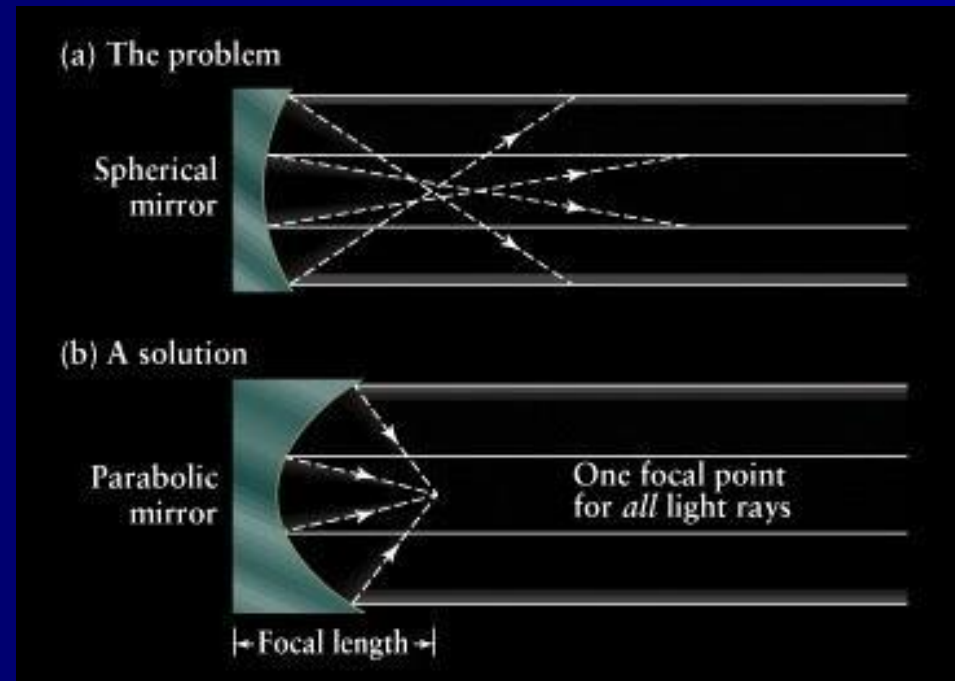


Alt-Azimuth Mount



Dobsonian Mount

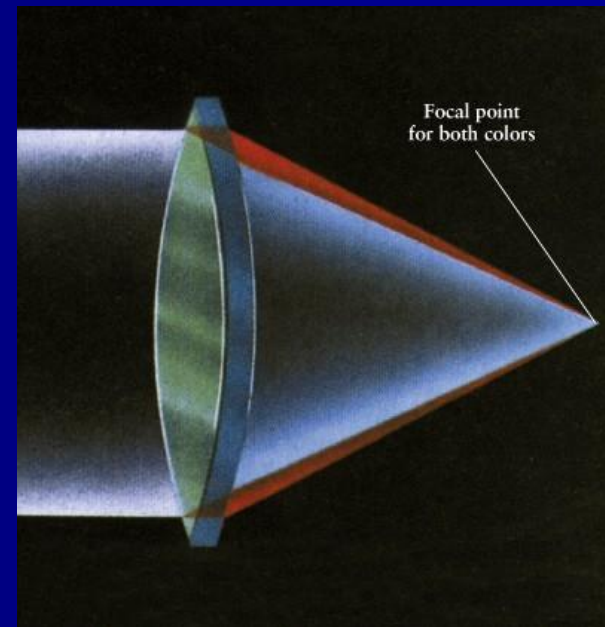
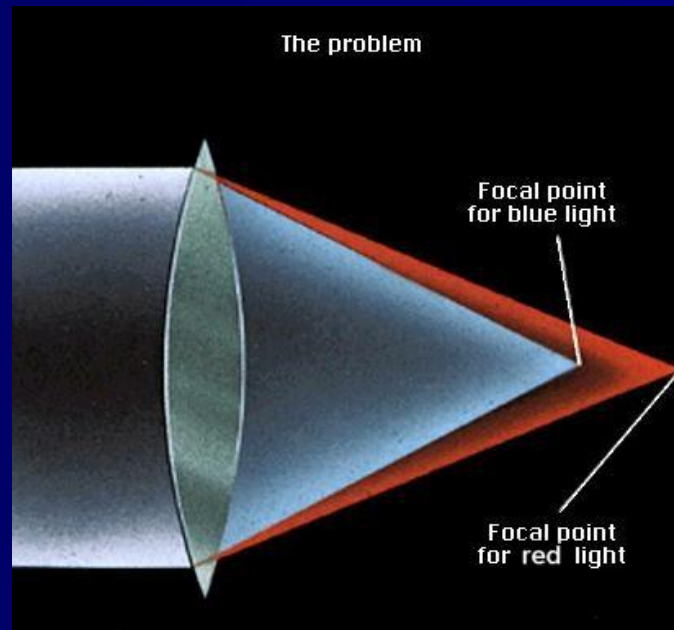
# Optical Aberrations



**Spherical Aberration**

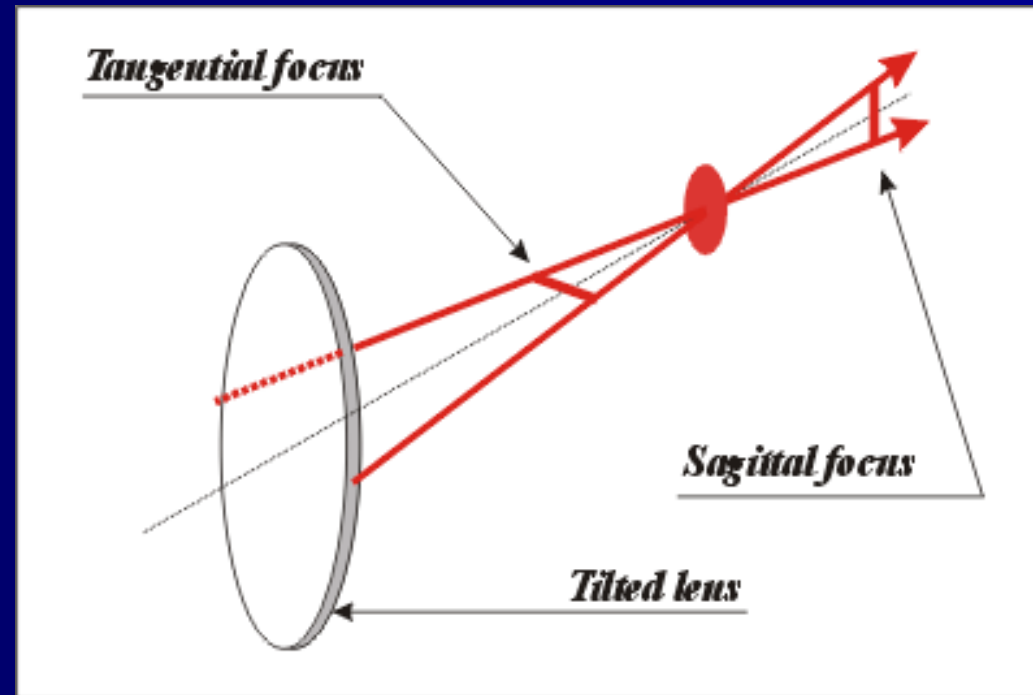


# Optical Aberrations



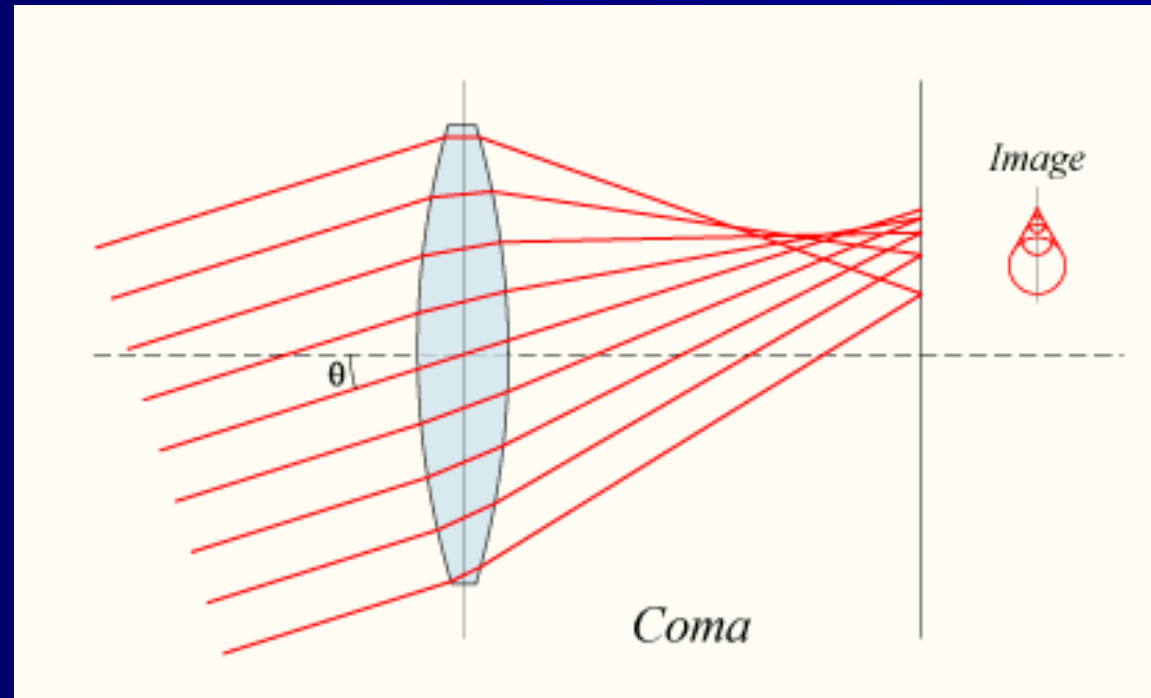
**Chromatic Aberration**

# Optical Aberrations



**Astigmatism**

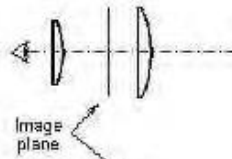
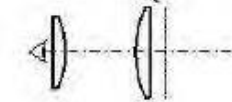
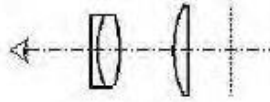
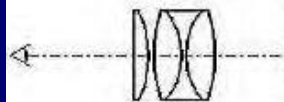
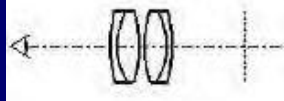

# Optical Aberrations



Coma

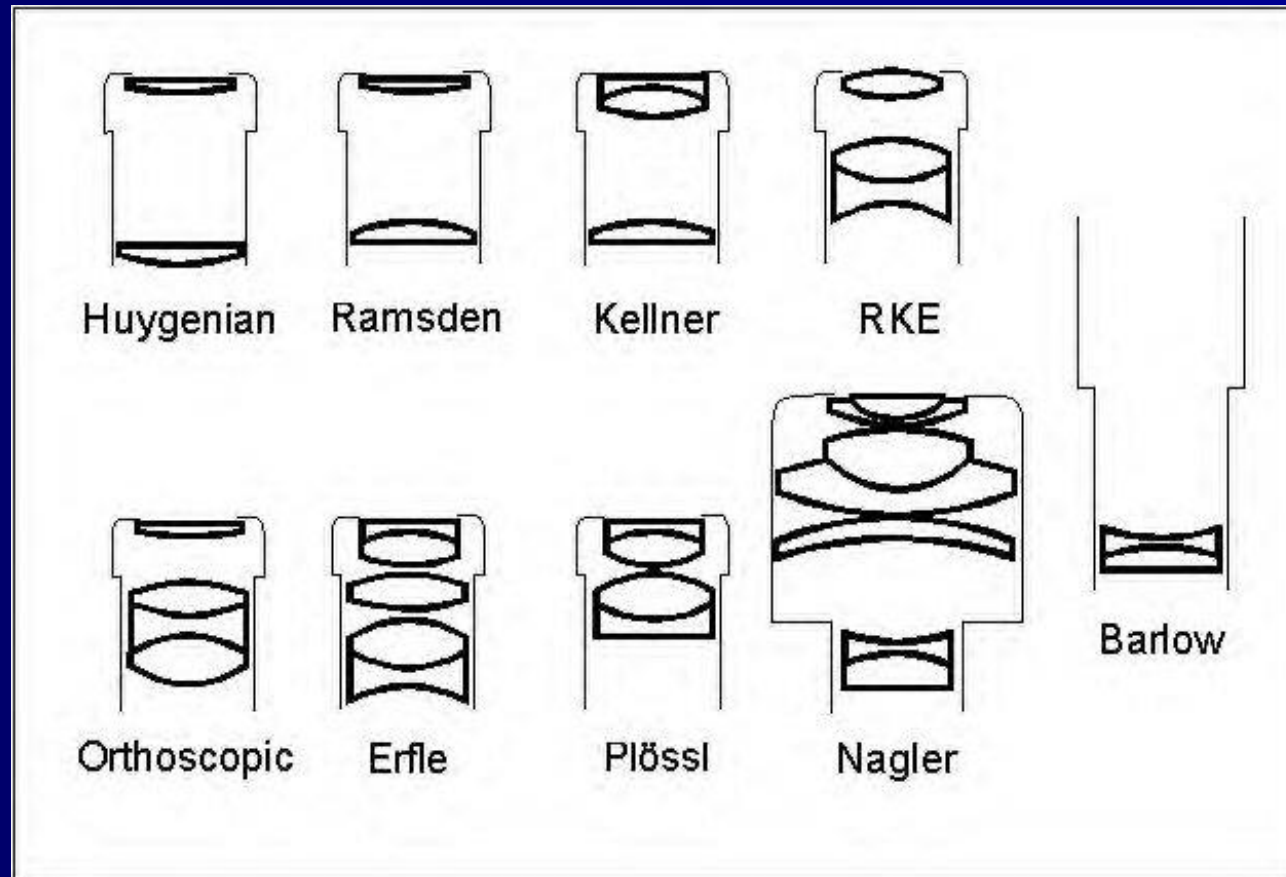
# Eyepieces (oculars)



	<b>HUYGENS EYEPIECE</b> This eyepiece reduces chromatic aberration to a minimum; with proper lens design, it can also reduce spherical aberration. In this eyepiece, the image plane is located between the field lens and the eye lens, the apparent field aperture is about 30°. It is a simple, widely used, and cheap eyepiece.
	<b>RAMSDEN EYEPIECE</b> The Ramsden eyepiece was invented 80 years after the Huygens eyepiece. Although it is the simplest eyepiece, it suffers from some limitations. The image must be very close to the field lens, so if the lens gets dusty, the dust will be quite evident. The eye distance is very short, so the lens may mist up or get dirty from contact with eyelashes. The field of view is 30-40°. In this eyepiece, as well as all those below, the image produced by the objective lens has to be focused in front of the field lens.
	<b>KELLNER EYEPIECE</b> The Kellner eyepiece uses an achromatic doublet to control chromatic aberration; spherical aberration is also minimal, and the eye relief is good. Field of view is 35-50°.
	<b>ABBE OR ORTHOSCOPIC EYEPIECE</b> Aberrations, especially distortions, are very well controlled. Long eye distance allows use of eyeglasses during observations. Field of view is 40-50°. Abbe eyepieces are widely used in binoculars and telescopes.
	<b>PLOSSL EYEPIECE</b> Simple and symmetrical, it uses two achromatic doublets. Performance is similar to the Abbe or orthoscopic eyepiece. Field of view is about 50°.
	<b>ERFLE EYEPIECE</b> Well corrected for aberrations, it produces a very wide field of about 70°. It is widely used in binoculars and other panoramic instruments.



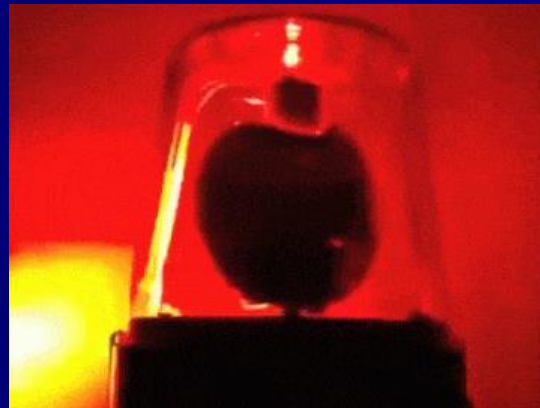
# Eyepieces (oculars)



# Warning!



# Math Alert



# Telescope Resolution



Dawes Limit or “D” apparent size of smallest object clearly resolved

$$D = \frac{4.56}{\text{aperature of telelscop e in inches}}$$

$$D = \frac{4.56}{6} = 0.76 \text{ arcsec}$$

# Magnification



$$M = \frac{\text{telescope focal length}}{\text{eyepiece focal length}}$$

**UNITS MUST BE THE SAME!!** 1in = 25.4mm

$$M = \frac{2000\text{mm}}{20\text{mm}} = 100X$$

# Limiting Magnitude



Faintest magnitude that can theoretically be seen.

$$M_l = 4.4 + 4.5 \log A$$

Where: **A** = the aperture in mm

# Telescope Types



## Refractors

### Pros

- Closed tube construction
- Unobstructed aperture
- High contrast images
- Holds collimation well
- Low maintenance
- No Coma

### Cons

- Not portable above 4"
- Needs heavier mount
- Chromatic aberration
- Difficult to manufacture
- High initial costs
- Eyepiece positioning



# Telescope Types



## Reflectors

### Pros

- Totally Achromatic
- Ease of portability
- Ease of manufacture
- Wide field of view
- Lower initial costs
- More aperture per \$

### Cons

- Central Obstruction
- Optical coatings deteriorate
- Easily misaligned
- More complex optical mounting
- Field restricted by coma
- Tube air currents

# Achromat vs. Apochromats

FYI Only



## Achromat

- Two widely spaced colors brought together
- Generally crown & flint doublet sometimes coated
- About 5% chromatic aberration
- Reasonable sharp and bright images on better scopes
- More expensive than reflector
  
- Good 4" Meade is \$900 - \$1200

## Apochromat

- Three widely spaced colors brought together
- Doublet with exotic materials such as calcium fluoride
- Color error reduced by 10%
- Slightly sharper & brighter (claimed by some)
- Double or more the cost of comparable achromat
- Meade 4" Apo is \$2000 - \$2500
- Takahashi 4" \$6,000 - \$8000

# Telescopes



**Thank You**

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Facebook Group**

